

Kerr-Newman spacetime

This worksheet demonstrates a few capabilities of [SageManifolds](#) (version 1.0, as included in SageMath 7.5) in computations regarding Kerr-Newman spacetime.

Click [here](#) to download the worksheet file (ipynb format). To run it, you must start SageMath within the Jupyter notebook, via the command `sage -n jupyter`

NB: a version of SageMath at least equal to 7.5 is required to run this worksheet:

```
In [1]: version()
```

```
Out[1]: 'SageMath version 7.5, Release Date: 2017-01-11'
```

First we set up the notebook to display mathematical objects using LaTeX rendering:

```
In [2]: %display latex
```

We also define a viewer for 3D plots (use 'threejs' or 'jmol' for interactive 3D graphics):

```
In [3]: viewer3D = 'jmol' # must be 'threejs', 'jmol', 'tachyon' or None (default)
```

Since some computations are quite long, we ask for running them in parallel on 8 cores:

```
In [4]: Parallelism().set(nproc=8)
```

Spacetime manifold

We declare the Kerr-Newman spacetime as a 4-dimensional differentiable manifold:

```
In [5]: M = Manifold(4, 'M', r'\mathcal{M}')
```

Let us use the standard **Boyer-Lindquist coordinates** on it, by first introducing the part \mathcal{M}_0 covered by these coordinates

```
In [6]: M0 = M.open_subset('M0', r'\mathcal{M}_0')
# BL = Boyer-Lindquist
BL.<t,r,th,ph> = M0.chart(r't r:(0,+oo) th:(0,pi):\theta ph:(0,2*pi):\phi')
print(BL) ; BL
```

```
Chart (M0, (t, r, th, ph))
```

```
Out[6]: ( $\mathcal{M}_0, (t, r, \theta, \phi)$ )
```

Metric tensor

The 3 parameters m , a and q of the Kerr-Newman spacetime are declared as symbolic variables:

In [7]: `var('m a q')`

Out[7]: (m, a, q)

Let us introduce the spacetime metric:

In [8]: `g = M.lorentzian_metric('g')`

The metric is defined by its components in the coordinate frame associated with Boyer-Lindquist coordinates, which is the current manifold's default frame:

```
In [9]: rho2 = r^2 + (a*cos(th))^2
Delta = r^2 - 2*m*r + a^2 + q^2
g[0,0] = -1 + (2*m*r - q^2)/rho2
g[0,3] = -a*sin(th)^2*(2*m*r - q^2)/rho2
g[1,1], g[2,2] = rho2/Delta, rho2
g[3,3] = (r^2 + a^2 + (2*m*r - q^2)*(a*sin(th))^2/rho2)*sin(th)^2
g.display()
```

Out[9]:

$$g = \left(-\frac{q^2 - 2mr}{a^2 \cos(\theta)^2 + r^2} - 1 \right) dt \otimes dt + \left(\frac{(q^2 - 2mr)a \sin(\theta)^2}{a^2 \cos(\theta)^2 + r^2} \right) dt \otimes d\phi$$

$$+ \left(\frac{a^2 \cos(\theta)^2 + r^2}{a^2 + q^2 - 2mr + r^2} \right) dr \otimes dr + (a^2 \cos(\theta)^2 + r^2) d\theta \otimes d\theta$$

$$+ \left(\frac{(q^2 - 2mr)a \sin(\theta)^2}{a^2 \cos(\theta)^2 + r^2} \right) d\phi \otimes dt$$

$$- \left(\frac{(q^2 - 2mr)a^2 \sin(\theta)^2}{a^2 \cos(\theta)^2 + r^2} - a^2 - r^2 \right) \sin(\theta)^2 d\phi \otimes d\phi$$

The list of the non-vanishing components:

In [10]: `g.display_comp()`

Out[10]:

$$g_{tt} = -\frac{q^2 - 2mr}{a^2 \cos(\theta)^2 + r^2} - 1$$

$$g_{t\phi} = \frac{(q^2 - 2mr)a \sin(\theta)^2}{a^2 \cos(\theta)^2 + r^2}$$

$$g_{rr} = \frac{a^2 \cos(\theta)^2 + r^2}{a^2 + q^2 - 2mr + r^2}$$

$$g_{\theta\theta} = a^2 \cos(\theta)^2 + r^2$$

$$g_{\phi t} = \frac{(q^2 - 2mr)a \sin(\theta)^2}{a^2 \cos(\theta)^2 + r^2}$$

$$g_{\phi\phi} = -\left(\frac{(q^2 - 2mr)a^2 \sin(\theta)^2}{a^2 \cos(\theta)^2 + r^2} - a^2 - r^2 \right) \sin(\theta)^2$$

The component g^{tt} of the inverse metric:

In [11]: `g.inverse()[0,0]`

Out[11]:

$$\frac{a^4 + 2a^2r^2 + r^4 - (a^4 + a^2q^2 - 2a^2mr + a^2r^2) \sin(\theta)^2}{2mr^3 - r^4 - (a^2 + q^2)r^2 - (a^4 + a^2q^2 - 2a^2mr + a^2r^2) \cos(\theta)^2}$$

The lapse function:

```
In [12]: N = 1/sqrt(-(g.inverse()[[0,0]])) ; N
```

Out[12]: Scalar field on the Open subset M0 of the 4-dimensional differentiable manifold M

```
In [13]: N.display()
```

Out[13]: $\mathcal{M}_0 \longrightarrow \mathbb{R}$
 $(t, r, \theta, \phi) \longmapsto \frac{\sqrt{a^2 \cos(\theta)^2 + r^2} \sqrt{a^2 + q^2 - 2mr + r^2}}{\sqrt{a^4 + 2a^2r^2 + r^4 - (a^4 + a^2q^2 - 2a^2mr + a^2r^2) \sin(\theta)^2}}$

Electromagnetic field tensor

Let us first introduce the 1-form basis associated with Boyer-Lindquist coordinates:

```
In [14]: dBL = BL.coframe() ; dBL
```

Out[14]: $(\mathcal{M}_0, (dt, dr, d\theta, d\phi))$

The electromagnetic field tensor F is formed as [cf. e.g. Eq. (33.5) of Misner, Thorne & Wheeler (1973)]

```
In [15]: F = M.diff_form(2, name='F')
F.set_restriction( q/rho2^2 * (r^2 - a^2*cos(th)^2) * dBL[1].wedge( dBL[0]
- a*sin(th)^2 * dBL[3] ) + \
2*q/rho2^2 * a*r*cos(th)*sin(th) * dBL[2].wedge( (r^2+a^2) * dBL[3] -
a* dBL[0] ) )
F.display()
```

Out[15]:
$$F = \left(\frac{a^2 q \cos(\theta)^2 - qr^2}{a^4 \cos(\theta)^4 + 2a^2 r^2 \cos(\theta)^2 + r^4} \right) dt \wedge dr$$

$$+ \left(\frac{2a^2 qr \cos(\theta) \sin(\theta)}{a^4 \cos(\theta)^4 + 2a^2 r^2 \cos(\theta)^2 + r^4} \right) dt \wedge d\theta$$

$$+ \left(\frac{(a^3 q \cos(\theta)^2 - aqr^2) \sin(\theta)^2}{a^4 \cos(\theta)^4 + 2a^2 r^2 \cos(\theta)^2 + r^4} \right) dr \wedge d\phi$$

$$+ \left(\frac{2(a^3 qr + aqr^3) \cos(\theta) \sin(\theta)}{a^4 \cos(\theta)^4 + 2a^2 r^2 \cos(\theta)^2 + r^4} \right) d\theta \wedge d\phi$$

The list of non-vanishing components:

```
In [16]: F.display_comp()
```

$$\begin{aligned}
 \text{Out[16]: } F_{tr} &= \frac{a^2 q \cos(\theta)^2 - qr^2}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \\
 F_{t\theta} &= \frac{2 a^2 q r \cos(\theta) \sin(\theta)}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \\
 F_{rt} &= -\frac{a^2 q \cos(\theta)^2 - qr^2}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \\
 F_{r\phi} &= \frac{(a^3 q \cos(\theta)^2 - aqr^2) \sin(\theta)^2}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \\
 F_{\theta t} &= -\frac{2 a^2 q r \cos(\theta) \sin(\theta)}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \\
 F_{\theta\phi} &= \frac{2 (a^3 q r + aqr^3) \cos(\theta) \sin(\theta)}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \\
 F_{\phi r} &= -\frac{(a^3 q \cos(\theta)^2 - aqr^2) \sin(\theta)^2}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \\
 F_{\phi\theta} &= -\frac{2 (a^3 q r + aqr^3) \cos(\theta) \sin(\theta)}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4}
 \end{aligned}$$

The Hodge dual of F :

```
In [17]: star_F = F.hodge_dual(g) ; star_F.display()
```

$$\begin{aligned}
 \text{Out[17]: } \star F &= \left(\frac{2 aqr \cos(\theta)}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \right) dt \wedge dr \\
 &+ \left(-\frac{(a^3 q \cos(\theta)^2 - aqr^2) \sin(\theta)}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \right) dt \wedge d\theta \\
 &+ \left(-\frac{2 (a^4 qr \cos(\theta) \sin(\theta)^4 - (a^4 qr + a^2 qr^3) \cos(\theta) \sin(\theta)^2)}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \right) dr \wedge d\phi \\
 &+ \left(\frac{(a^4 q + a^2 qr^2) \sin(\theta)^3 - (a^4 q - qr^4) \sin(\theta)}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \right) d\theta \wedge d\phi
 \end{aligned}$$

Maxwell equations

Let us check that F obeys the two (source-free) Maxwell equations:

```
In [18]: F.exterior_derivative().display()
```

$$\text{Out[18]: } dF = 0$$

```
In [19]: star_F.exterior_derivative().display()
```

$$\text{Out[19]: } d \star F = 0$$

Levi-Civita Connection

The Levi-Civita connection ∇ associated with g :

```
In [20]: nab = g.connection() ; print(nab)
```

Levi-Civita connection `nabla_g` associated with the Lorentzian metric g on the 4-dimensional differentiable manifold M

Let us verify that the covariant derivative of g with respect to ∇ vanishes identically:

```
In [21]: nab(g) == 0
```

```
Out[21]: True
```

Another view of the above property:

```
In [22]: nab(g).display()
```

```
Out[22]:  $\nabla_g g = 0$ 
```

The nonzero Christoffel symbols (skipping those that can be deduced by symmetry of the last two indices):

In [23]: `g.christoffel_symbols_display()`

Out[23]:

$$\begin{aligned} \Gamma^t{}_{tr} &= \frac{a^4 m + a^2 q^2 r + q^2 r^3 - m r^4 - (a^4 m + a^2 m r^2) \sin(\theta)^2}{2 m r^5 - r^6 - (a^2 + q^2) r^4 - (a^6 + a^4 q^2 - 2 a^4 m r + a^4 r^2) \cos(\theta)^4 + 2 (2 a^2 m r^3 - a^2 r^4 - (a^4 + a^2 q^2) r^2) \cos(\theta)} \\ \Gamma^t{}_{t\theta} &= \frac{(a^2 q^2 - 2 a^2 m r) \cos(\theta) \sin(\theta)}{a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4} \\ \Gamma^t{}_{r\phi} &= -\frac{a^3 q^2 r - a^3 m r^2 + 2 a q^2 r^3 - 3 a m r^4 - (a^5 m + a^3 q^2 r - a^3 m r^2) \cos(\theta)^4 + (a^5 m - 2 a q^2 r^3 + 3 a m r^4) \cos(\theta)}{2 m r^5 - r^6 - (a^2 + q^2) r^4 - (a^6 + a^4 q^2 - 2 a^4 m r + a^4 r^2) \cos(\theta)^4 + 2 (2 a^2 m r^3 - a^2 r^4 - (a^4 + a^2 q^2) r^2) \cos(\theta)} \\ \Gamma^t{}_{\theta\phi} &= \frac{(a^5 q^2 - 2 a^5 m r) \cos(\theta) \sin(\theta)^5 - (a^5 q^2 - 2 a^5 m r + a^3 q^2 r^2 - 2 a^3 m r^3) \cos(\theta) \sin(\theta)^3}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \\ \Gamma^r{}_{tt} &= \frac{m r^4 - (2 m^2 + q^2) r^3 + (a^2 m + 3 m q^2) r^2 - (a^4 m + a^2 m q^2 - 2 a^2 m^2 r + a^2 m r^2) \cos(\theta)^2 - (a^2 q^2 + q^4) r}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \\ \Gamma^r{}_{t\phi} &= -\frac{(a m r^4 - (2 a m^2 + a q^2) r^3 + (a^3 m + 3 a m q^2) r^2 - (a^5 m + a^3 m q^2 - 2 a^3 m^2 r + a^3 m r^2) \cos(\theta)^2 - (a^3 q^2 + a q^4)}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \\ \Gamma^r{}_{rr} &= -\frac{a^2 m + q^2 r - m r^2 - (a^2 m - a^2 r) \sin(\theta)^2}{2 m r^3 - r^4 - (a^2 + q^2) r^2 - (a^4 + a^2 q^2 - 2 a^2 m r + a^2 r^2) \cos(\theta)^2} \\ \Gamma^r{}_{r\theta} &= -\frac{a^2 \cos(\theta) \sin(\theta)}{a^2 \cos(\theta)^2 + r^2} \\ \Gamma^r{}_{\theta\theta} &= \frac{2 m r^2 - r^3 - (a^2 + q^2) r}{a^2 \cos(\theta)^2 + r^2} \\ \Gamma^r{}_{\phi\phi} &= \frac{(a^2 m r^4 - (2 a^2 m^2 + a^2 q^2) r^3 + (a^4 m + 3 a^2 m q^2) r^2 - (a^6 m + a^4 m q^2 - 2 a^4 m^2 r + a^4 m r^2) \cos(\theta)^2 - (a^4 q^2 + a^2 \theta)^4 + (2 m r^6 - r^7 - (a^2 + q^2) r^5 + (2 a^4 m r^2 - a^4 r^3 - (a^6 + a^4 q^2) r) \cos(\theta)^4 + 2 (2 a^2 m r^4 - a^2 r^5 - (a^4 + a^2 q^2) \theta)^2)}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \\ \Gamma^\theta{}_{tt} &= \frac{(a^2 q^2 - 2 a^2 m r) \cos(\theta) \sin(\theta)}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \\ \Gamma^\theta{}_{t\phi} &= -\frac{(a^3 q^2 - 2 a^3 m r + a q^2 r^2 - 2 a m r^3) \cos(\theta) \sin(\theta)}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \\ \Gamma^\theta{}_{rr} &= -\frac{a^2 \cos(\theta) \sin(\theta)}{2 m r^3 - r^4 - (a^2 + q^2) r^2 - (a^4 + a^2 q^2 - 2 a^2 m r + a^2 r^2) \cos(\theta)^2} \\ \Gamma^\theta{}_{r\theta} &= \frac{r}{a^2 \cos(\theta)^2 + r^2} \\ \Gamma^\theta{}_{\theta\theta} &= -\frac{a^2 \cos(\theta) \sin(\theta)}{a^2 \cos(\theta)^2 + r^2} \\ \Gamma^\theta{}_{\phi\phi} &= -\frac{((a^6 + a^4 q^2 - 2 a^4 m r + a^4 r^2) \cos(\theta)^5 - 2 (2 a^2 m r^3 - a^2 r^4 - (a^4 + a^2 q^2) r^2) \cos(\theta)^3 \sin(\theta) - (a^4 q^2 - 2 a^4 m r + 2 a^2 q^2 r^2 - 4 a^2 m r^3 - a^2 r^4 - r^6) \cos(\theta))}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \\ \Gamma^\phi{}_{tr} &= \frac{a^3 m \cos(\theta)^2 + a q^2 r - a m r^2}{2 m r^5 - r^6 - (a^2 + q^2) r^4 - (a^6 + a^4 q^2 - 2 a^4 m r + a^4 r^2) \cos(\theta)^4 + 2 (2 a^2 m r^3 - a^2 r^4 - (a^4 + a^2 q^2) r^2) \cos(\theta)} \\ \Gamma^\phi{}_{t\theta} &= \frac{(a q^2 - 2 a m r) \cos(\theta)}{(a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4) \sin(\theta)} \\ \Gamma^\phi{}_{r\phi} &= -\frac{a^2 q^2 r - a^2 m r^2 + q^2 r^3 - 2 m r^4 + r^5 - (a^4 m - a^4 r) \cos(\theta)^4 + (a^4 m - a^2 m r^2 + 2 a^2 r^3) \cos(\theta)^2}{2 m r^5 - r^6 - (a^2 + q^2) r^4 - (a^6 + a^4 q^2 - 2 a^4 m r + a^4 r^2) \cos(\theta)^4 + 2 (2 a^2 m r^3 - a^2 r^4 - (a^4 + a^2 q^2) r^2) \cos(\theta)} \\ \Gamma^\phi{}_{\theta\phi} &= \frac{a^4 \cos(\theta) \sin(\theta)^4 - (2 a^4 + a^2 q^2 - 2 a^2 m r + 2 a^2 r^2) \cos(\theta) \sin(\theta)^2 + (a^4 + 2 a^2 r^2 + r^4) \cos(\theta)}{(a^4 \cos(\theta)^4 + 2 a^2 r^2 \cos(\theta)^2 + r^4) \sin(\theta)} \end{aligned}$$

Killing vectors

The default vector frame on the spacetime manifold is the coordinate basis associated with Boyer-Lindquist coordinates:

```
In [24]: M.default_frame() is BL.frame()
```

```
Out[24]: True
```

```
In [25]: BL.frame()
```

```
Out[25]:  $\left(\mathcal{M}_0, \left(\frac{\partial}{\partial t}, \frac{\partial}{\partial r}, \frac{\partial}{\partial \theta}, \frac{\partial}{\partial \phi}\right)\right)$ 
```

Let us consider the first vector field of this frame:

```
In [26]: xi = BL.frame()[0] ; xi
```

```
Out[26]:  $\frac{\partial}{\partial t}$ 
```

```
In [27]: print(xi)
```

Vector field d/dt on the Open subset M0 of the 4-dimensional differentiable manifold M

The 1-form associated to it by metric duality is

```
In [28]: xi_form = xi.down(g) ; xi_form.display()
```

```
Out[28]:  $\left(-\frac{a^2 \cos(\theta)^2 + q^2 - 2mr + r^2}{a^2 \cos(\theta)^2 + r^2}\right) dt + \left(\frac{(aq^2 - 2amr) \sin(\theta)^2}{a^2 \cos(\theta)^2 + r^2}\right) d\phi$ 
```

Its covariant derivative is

In [29]: `nab_xi = nab(xi_form) ; print(nab_xi) ; nab_xi.display()`

Tensor field of type (0,2) on the Open subset M0 of the 4-dimensional differentiable manifold M

$$\begin{aligned}
 \text{Out[29]: } & \left(\frac{mr^4 - (2m^2 + q^2)r^3 + (a^2m + 3mq^2)r^2}{2mr^5 - r^6 - (a^2 + q^2)r^4 - (a^6 + a^4q^2 - 2a^4mr + a^4r^2)\cos(\theta)^4 + 2} \right) dt \\
 & \otimes dr + \left(-\frac{(a^2q^2 - 2a^2mr)\cos(\theta)\sin(\theta)}{a^4\cos(\theta)^4 + 2a^2r^2\cos(\theta)^2 + r^4} \right) dt \otimes d\theta \\
 & + \left(-\frac{mr^4 - (2m^2 + q^2)r^3 + (a^2m + 3mq^2)r^2}{2mr^5 - r^6 - (a^2 + q^2)r^4 - (a^6 + a^4q^2 - 2a^4mr + a^4r^2)\cos(\theta)^4 + 2} \right) dr \\
 & \otimes dt + \left(-\frac{a^3m\cos(\theta)^4 - aq^2r + amr^2 - (a^3m - aq^2r + amr^2)\cos(\theta)^2}{a^4\cos(\theta)^4 + 2a^2r^2\cos(\theta)^2 + r^4} \right) dr \\
 & \otimes d\phi + \left(\frac{(a^2q^2 - 2a^2mr)\cos(\theta)\sin(\theta)}{a^4\cos(\theta)^4 + 2a^2r^2\cos(\theta)^2 + r^4} \right) d\theta \otimes dt \\
 & + \left(-\frac{(a^3q^2 - 2a^3mr + aq^2r^2 - 2amr^3)\cos(\theta)\sin(\theta)}{a^4\cos(\theta)^4 + 2a^2r^2\cos(\theta)^2 + r^4} \right) d\theta \otimes d\phi \\
 & + \left(\frac{a^3m\sin(\theta)^4 - (a^3m + aq^2r - amr^2)\sin(\theta)^2}{a^4\cos(\theta)^4 + 2a^2r^2\cos(\theta)^2 + r^4} \right) d\phi \otimes dr \\
 & + \left(\frac{(a^3q^2 - 2a^3mr + aq^2r^2 - 2amr^3)\cos(\theta)\sin(\theta)}{a^4\cos(\theta)^4 + 2a^2r^2\cos(\theta)^2 + r^4} \right) d\phi \otimes d\theta
 \end{aligned}$$

Let us check that the vector field $\xi = \frac{\partial}{\partial r}$ obeys Killing equation:

In [30]: `nab_xi.symmetrize() == 0`

Out[30]: True

Similarly, let us check that $\chi := \frac{\partial}{\partial \phi}$ is a Killing vector:

In [31]: `chi = BL.frame()[3] ; chi`

Out[31]: $\frac{\partial}{\partial \phi}$

In [32]: `nab(chi.down(g)).symmetrize() == 0`

Out[32]: True

Another way to check that ξ and χ are Killing vectors is the vanishing of the Lie derivative of the metric tensor along them:


```
In [33]: g.lie_derivative(xi) == 0
```

```
Out[33]: True
```

```
In [34]: g.lie_derivative(chi) == 0
```

```
Out[34]: True
```

Curvature

The Ricci tensor associated with g :

```
In [35]: Ric = g.ricci() ; print(Ric)
```

Field of symmetric bilinear forms Ric(g) on the 4-dimensional differentiable manifold M

```
In [36]: Ric.display()
```

```
Out[36]:
```

$$\begin{aligned}
 \text{Ric}(g) = & \left(-\frac{a^2 q^2 \cos(\theta)^2 - 2 a^2 q^2 - q^4 + 2 m q^2 r - q^2 r^2}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \right) dt \otimes dt \\
 & + \left(\frac{2 a^3 q^2 + a q^4 - 2 a m q^2 r + 2 a q^2 r^2 - (2 a^3 q^2 + a q^4 - 2 a m q^2 r + 2 a q^2 r^2) \cos(\theta)^2}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \right) \\
 & \otimes d\phi + \left(\frac{q^2}{2 m r^3 - r^4 - (a^2 + q^2) r^2 - (a^4 + a^2 q^2 - 2 a^2 m r + a^2 r^2) \cos(\theta)^2} \right) dr \\
 & \otimes dr + \left(\frac{q^2}{a^2 \cos(\theta)^2 + r^2} \right) d\theta \otimes d\theta \\
 & + \left(\frac{2 a^3 q^2 + a q^4 - 2 a m q^2 r + 2 a q^2 r^2 - (2 a^3 q^2 + a q^4 - 2 a m q^2 r + 2 a q^2 r^2) \cos(\theta)^2}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} \right) \\
 & \otimes dt + \left(\frac{(a^6 q^2 + a^4 q^4 - 2 a^4 m q^2 r + a^4 q^2 r^2) \sin(\theta)^6 - (a^4 q^4 - 2 a^4 m q^2 r + a^2 q^4 r^2 - 2 a^2 m q^2 r^3) \sin(\theta)^4 - (a^6 q^2 + 3 a^4 q^2 r^2 + 3 a^2 q^2 r^4 + q^2 r^6) \sin(\theta)^2}{a^8 \cos(\theta)^8 + 4 a^6 r^2 \cos(\theta)^6 + 6 a^4 r^4 \cos(\theta)^4 + 4 a^2 r^6 \cos(\theta)^2 + r^8} \right) d\phi \\
 & \otimes d\phi
 \end{aligned}$$

In [37]: Ric[:]

$$\text{Out[37]: } \begin{pmatrix} -\frac{a^2 q^2 \cos(\theta)^2 - 2 a^2 q^2 - q^4 + 2 m q^2 r - q^2 r^2}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} & 0 & \frac{q^2}{2 m r^3 - r^4 - (a^2 + q^2) r^2 - (a^4 + a^2 q^2 -} \\ & 0 & \\ -\frac{2 a^3 q^2 + a q^4 - 2 a m q^2 r + 2 a q^2 r^2 - (2 a^3 q^2 + a q^4 - 2 a m q^2 r + 2 a q^2 r^2) \cos(\theta)^2}{a^6 \cos(\theta)^6 + 3 a^4 r^2 \cos(\theta)^4 + 3 a^2 r^4 \cos(\theta)^2 + r^6} & & \end{pmatrix}$$

Let us check that in the Kerr case, i.e. when $q = 0$, the Ricci tensor is zero:

In [38]: Ric[:].subs(q=0)

$$\text{Out[38]: } \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$

The Riemann curvature tensor associated with g :

In [39]: R = g.riemann() ; print(R)

Tensor field Riem(g) of type (1,3) on the 4-dimensional differentiable manifold M

The component R^0_{101} of the Riemann tensor is

In [40]: R[0,1,0,1]

$$\text{Out[40]: } \frac{4 a^2 q^2 r^2 - 3 a^2 m r^3 + 3 q^2 r^4 - 2 m r^5 + (a^4 q^2 - 3 a^4 m r) \cos(\theta)^4 - (2 a^4 q^2 - 9 a^4 m r + 2 a^2 q^2 r^2 - 7 a^2 m r^3) \cos(\theta)^2}{2 m r^7 - r^8 - (a^2 + q^2) r^6 - (a^8 + a^6 q^2 - 2 a^6 m r + a^6 r^2) \cos(\theta)^6 + 3 (2 a^4 m r^3 - a^4 r^4 - (a^6 + a^4 q^2) r^2) \cos(\theta)^4 + 3 (2 a^2 m r^5 - a^2 r^6 - (a^4 + a^2 q^2) r^4) \cos(\theta)^2}$$

The expression in the uncharged limit (Kerr spacetime) is

In [41]: R[0,1,0,1].expr().subs(q=0).simplify_rational()

$$\text{Out[41]: } \frac{3 a^4 m r \cos(\theta)^4 + 3 a^2 m r^3 + 2 m r^5 - (9 a^4 m r + 7 a^2 m r^3) \cos(\theta)^2}{a^2 r^6 - 2 m r^7 + r^8 + (a^8 - 2 a^6 m r + a^6 r^2) \cos(\theta)^6 + 3 (a^6 r^2 - 2 a^4 m r^3 + a^4 r^4) \cos(\theta)^4 + 3 (a^4 r^4 - 2 a^2 m r^5 + a^2 r^6) \cos(\theta)^2}$$

while in the non-rotating limit (Reissner-Nordström spacetime), it is

```
In [42]: R[0,1,0,1].expr().subs(a=0).simplify_rational()
```

Out[42]:
$$-\frac{3q^2 - 2mr}{q^2r^2 - 2mr^3 + r^4}$$

In the Schwarzschild limit, it reduces to

```
In [43]: R[0,1,0,1].expr().subs(a=0, q=0).simplify_rational()
```

Out[43]:
$$-\frac{2m}{2mr^2 - r^3}$$

Obviously, it vanishes in the flat space limit:

```
In [44]: R[0,1,0,1].expr().subs(m=0, a=0, q=0)
```

Out[44]: 0

Bianchi identity

Let us check the Bianchi identity $\nabla_p R^i_{jkl} + \nabla_k R^i_{jlp} + \nabla_l R^i_{jpk} = 0$:

```
In [45]: DR = nab(R) ; print(DR) #long (takes a while)
```

Tensor field `nabla_g(Riem(g))` of type (1,4) on the 4-dimensional differentiable manifold `M`


```
In [49]: g.ricci_scalar().display()
```

```
Out[49]: r(g) :      M      → ℝ
          on M0 : (t, r, θ, φ) ↦ 0
```

Einstein equation

The Einstein tensor is

```
In [50]: G = Ric - 1/2*g.ricci_scalar()*g ; print(G)
```

```
Field of symmetric bilinear forms +Ric(g) on the 4-dimensional differen
tible manifold M
```

Since the Ricci scalar is zero, the Einstein tensor reduces to the Ricci tensor:

```
In [51]: G == Ric
```

```
Out[51]: True
```

The invariant $F_{ab}F^{ab}$ of the electromagnetic field:

```
In [52]: Fuu = F.up(g)
         F2 = F['_ab']*Fuu['^ab'] ; print(F2)
```

```
Scalar field on the 4-dimensional differentiable manifold M
```

```
In [53]: F2.display()
```

```
Out[53]:      M      → ℝ
          on M0 : (t, r, θ, φ) ↦ 
$$-\frac{2(a^4q^2\cos(\theta)^4 - 6a^2q^2r^2\cos(\theta)^2 + q^2r^4)}{a^8\cos(\theta)^8 + 4a^6r^2\cos(\theta)^6 + 6a^4r^4\cos(\theta)^4 + 4a^2r^6\cos(\theta)^2 + r^8}$$

```

The energy-momentum tensor of the electromagnetic field:

```
In [54]: Fud = F.up(g, θ)
         T = 1/(4*pi)*( F['_k_']*Fud['^k_'] - 1/4*F2 * g ); print(T)
```

```
Tensor field of type (0,2) on the 4-dimensional differentiable manifold
M
```

```
In [55]: T[:]
```

```
Out[55]: 
$$\begin{pmatrix} -\frac{a^2q^2\cos(\theta)^2 - 2a^2q^2 - q^4 + 2mq^2r - q^2r^2}{8(\pi a^6\cos(\theta)^6 + 3\pi a^4r^2\cos(\theta)^4 + 3\pi a^2r^4\cos(\theta)^2 + \pi r^6)} & 0 & 0 \\ & 0 & \frac{q^2}{8(2\pi mr^3 - \pi r^4 - (\pi a^2 + \pi q^2)r^2 - (\pi a^4 + \pi a^2q^2 - 2\pi a^2r^2))} \\ & & 0 \\ & & & -\frac{(2a^3q^2 + aq^4 - 2amq^2r + 2aq^2r^2)\sin(\theta)^2}{8(\pi a^6\cos(\theta)^6 + 3\pi a^4r^2\cos(\theta)^4 + 3\pi a^2r^4\cos(\theta)^2 + \pi r^6)} \end{pmatrix}$$

```

Check of the Einstein equation:

In [56]: `G == 8*pi*T`

Out[56]: True

Kretschmann scalar

The tensor R^b , of components $R_{abcd} = g_{am}R^m{}_{bcd}$:

In [57]: `dR = R.down(g) ; print(dR)`

Tensor field of type (0,4) on the 4-dimensional differentiable manifold M

The tensor R^\sharp , of components $R^{abcd} = g^{bp}g^{cq}g^{dr}R^a{}_{pqr}$:

In [58]: `uR = R.up(g) ; print(uR)`

Tensor field of type (4,0) on the 4-dimensional differentiable manifold M

The Kretschmann scalar $K := R^{abcd}R_{abcd}$:

In [59]: `Kr_scalar = uR['^ijkl']*dR['_ijkl']
Kr_scalar.display()`

Out[59]:

$$\mathcal{M} \longrightarrow \mathbb{R}$$

$$\left(\begin{array}{l} 6 m^2 r^8 - 12 (m^3 + m q^2) r^7 + (6 a^2 m^2 + 30 m^2 q^2 + 7 q^4) r^6 - 6 (a^8 m^2 + a^6 m^2 q^2 + 6 a^2 m q^2 + 13 m q^4) r^5 + 7 (a^2 q^4 + q^6) r^4 \\ + (7 a^6 q^4 + 7 a^4 q^6 + 90 a^4 m^2 r^4 - 60 (3 a^4 m^3 + a^4 m q^2) r^3 + (90 a^6 m^2 + 210 a^4 (30 a^6 m q^2 + 37 a^4 m q^4) r) \\ (45 a^2 m^2 r^6 - 30 (3 a^2 m^3 + 2 a^2 m q^2) r^5 + (45 a^4 m^2 + 165 a^2 m^2 q^2 + 17 a^2 q^4) (a^4 q^4 + a^2 q^6) r^2) \\ (\theta^2) \end{array} \right)$$

$$\text{on } \mathcal{M}_0 : (t, r, \theta, \phi) \longmapsto \frac{\left(\begin{array}{l} 2 m r^{13} - r^{14} - (a^2 + q^2) r^{12} - (a^{14} + a^{12} q^2 - 2 a^{12} m r + a^{12} r^2) \cos(\theta)^{12} + 6 (2 (\theta)^{10} + 15 (2 a^8 m r^5 - a^8 r^6 - (a^{10} + a^8 q^2) r^4) \cos(\theta)^8 + 20 (2 a^6 m r^7 - a^6 r^8 - (2 a^4 m r^9 - a^4 r^{10} - (a^6 + a^4 q^2) r^8) \cos(\theta)^4 + 6 (2 a^2 m r^{11} - a^2 r^{12} - (a^4 + a^2 \end{array} \right)}{(a^2 \cos(\theta)^2 + r^2)^6}$$

A variant of this expression can be obtained by invoking the `factor()` method on the coordinate function representing the scalar field in the manifold's default chart:

In [60]: `Kr = Kr_scalar.coord_function()
Kr.factor()`

Out[60]:

$$\frac{8 (6 a^6 m^2 \cos(\theta)^6 - 7 a^4 q^4 \cos(\theta)^4 + 60 a^4 m q^2 r \cos(\theta)^4 - 90 a^4 m^2 r^2 \cos(\theta)^4 + 34 a^2 q^4 r^2 \cos(\theta)^2 - 120 a^2 m q^2 r^3 \cos(\theta)^2 + 90 a^2 m^2 r^4 \cos(\theta)^2 - 7 q^4 r^4 + 12 m q^2 r^5 - 6 m^2 r^6)}{(a^2 \cos(\theta)^2 + r^2)^6}$$

As a check, we can compare Kr to the formula given by R. Conn Henry, [Astrophys. J. 535, 350 \(2000\)](#):

```
In [61]: Kr == 8/(r^2+(a*cos(th))^2)^6 *(
          6*m^2*(r^6 - 15*r^4*(a*cos(th))^2 + 15*r^2*(a*cos(th))^4 - (a
          *cos(th))^6)
          - 12*m*q^2*r*(r^4 - 10*(a*r*cos(th))^2 + 5*(a*cos(th))^4)
          + q^4*(7*r^4 - 34*(a*r*cos(th))^2 + 7*(a*cos(th))^4 )
```

Out[61]: True

The Schwarzschild value of the Kretschmann scalar is recovered by setting $a = 0$ and $q = 0$:

```
In [62]: Kr.expr().subs(a=0, q=0)
```

Out[62]: $\frac{48 m^2}{r^6}$

Let us plot the Kretschmann scalar for $m = 1$, $a = 0.9$ and $q = 0.5$:

```
In [63]: K1 = Kr.expr().subs(m=1, a=0.9, q=0.5)
          plot3d(K1, (r,1,3), (th, 0, pi), viewer=viewer3D, axes_labels=['r', 'th
          eta', 'Kr'])
```

Out[63]:

