

3+1 Einstein equations in the $\delta = 2$ Tomimatsu-Sato spacetime

This worksheet demonstrates a few capabilities of [SageManifolds](#) (version 1.0, as included in SageMath 7.5) in computations regarding the 3+1 slicing of the $\delta = 2$ Tomimatsu-Sato spacetime.

Click [here](#) to download the worksheet file (ipynb format). To run it, you must start SageMath with the Jupyter notebook, via the command `sage -n jupyter`

NB: a version of SageMath at least equal to 7.5 is required to run this worksheet:

```
In [1]: version()
```

```
Out[1]: 'SageMath version 7.5.1, Release Date: 2017-01-15'
```

First we set up the notebook to display mathematical objects using LaTeX rendering:

```
In [2]: %display latex
```

Since some computations are quite long, we ask for running them in parallel on 8 cores:

```
In [3]: Parallelism().set(nproc=8)
```

Tomimatsu-Sato spacetime

The Tomimatsu-Sato solution is an exact stationary and axisymmetric solution of the vacuum Einstein equation, which is asymptotically flat and has a non-zero angular momentum. It has been found in 1972 by A. Tomimatsu and H. Sato [[Phys. Rev. Lett. 29, 1344 \(1972\)](#)], as a solution of the Ernst equation. It is actually the member $\delta = 2$ of a larger family of solutions parametrized by a positive integer δ and exhibited by Tomimatsu and Sato in 1973 [[Prog. Theor. Phys. 50, 95 \(1973\)](#)], the member $\delta = 1$ being nothing but the Kerr metric. We refer to [[Manko, Prog. Theor. Phys. 127, 1057 \(2012\)](#)] for a discussion of the properties of this solution.

Spacelike hypersurface

We consider some hypersurface Σ of a spacelike foliation $(\Sigma_t)_{t \in \mathbb{R}}$ of $\delta = 2$ Tomimatsu-Sato spacetime; we declare Σ_t as a 3-dimensional manifold:

```
In [4]: Sig = Manifold(3, 'Sigma', r'\Sigma', start_index=1)
```

On Σ , we consider the prolate spheroidal coordinates (x, y, ϕ) , with $x \in (1, +\infty)$, $y \in (-1, 1)$ and $\phi \in (0, 2\pi)$:

```
In [5]: X.<r,y,ph> = Sig.chart(r'x:(1,+oo) y:(-1,1) ph:(0,2*pi):\phi')
print(X) ; X
```

```
Chart (Sigma, (x, y, ph))
```

```
Out[5]: ( $\Sigma, (x, y, \phi)$ )
```

Riemannian metric on Σ

The Tomimatsu-Sato metric depends on three parameters: the integer δ , the real number $p \in [0, 1]$, and the total mass m :

```
In [6]: var('d, p, m')
        assume(m>0)
        assumptions()
```

```
Out[6]: [x is real, x > 1, y is real, y > (-1), y < 1, phi is real, phi > 0, phi < 2 pi,
         m > 0]
```

We set $\delta = 2$ and choose a specific value for $p = 1/5$:

```
In [7]: d = 2
        p = 1/5
```

Furthermore, without any loss of generality, we may set $m = 1$ (this simply fixes some length scale):

```
In [8]: m = 1
```

The parameter q is related to p by $p^2 + q^2 = 1$:

```
In [9]: q = sqrt(1-p^2)
```

Some shortcut notations:

```
In [10]: AA2 = (p^2*(x^2-1)^2+q^2*(1-y^2)^2) \
            - 4*p^2*q^2*(x^2-1)*(1-y^2)*(x^2-y^2)^2
        BB2 = (p^2*x^4+2*p*x^3-2*p*x+q^2*y^4-1)^2 \
            + 4*q^2*y^2*(p*x^3-p*x*y^2-y^2+1)^2
        CC2 = p^3*x*(1-x^2)*(2*(x^4-1)+(x^2+3)*(1-y^2)) \
            + p^2*(x^2-1)*((x^2-1)*(1-y^2)-4*x^2*(x^2-y^2)) \
            + q^2*(1-y^2)^3*(p*x+1)
```

The Riemannian metric γ induced by the spacetime metric g on Σ :

```
In [11]: gam = Sig.riemannian_metric('gam', latex_name=r'\gamma')
gam[1,1] = m^2*BB2/(p^2*d^2*(x^2-1)*(x^2-y^2)^3)
gam[2,2] = m^2*BB2/(p^2*d^2*(y^2-1)*(-x^2+y^2)^3)
gam[3,3] = - m^2*(y^2-1)*(p^2*BB2^2*(x^2-1)
+ 4*q^2*d^2*CC2^2*(y^2-1))/(AA2*BB2*d^2)
gam.display()
```

Out[11]:

$$\gamma = \left(\frac{96(x^3 - xy^2 - 5y^2 + 5)^2 y^2 + (x^4 + 24y^4 - 10x^3 - 10x - 25)^2}{100(x^2 - y^2)^3(x^2 - 1)} \otimes dx + \left(-\frac{96(x^3 - xy^2 - 5y^2 + 5)^2 y^2 + (x^4 + 24y^4 - 10x^3 - 10x - 25)^2}{100(x^2 - y^2)^3(y^2 - 1)} \otimes dy \right. \right.$$

$$\left. \left(\frac{(96(x^3 - xy^2 - 5y^2 + 5)^2 y^2 + (x^4 + 24y^4 + 10x^3 - 10x - 25)^2)(x^2 - 1) + 9600}{100} \right. \right.$$

$$\left. \frac{(24(y^2 - 1)^3(x + 5) + (2x^4 - (x^2 + 3)(y^2 - 1) - 2)(x^2 - 1) + 4(x^2 - y^2)x^2 + (x^2 - 1)(y^2 - 1))(x^2 - 1)}{100} \right)$$

$$\left(\frac{(96(x^2 - y^2)^2(x^2 - 1)(y^2 - 1) + ((x^2 - 1)^2 + 24(y^2 - 1)^2)^2)(96(x^3 - xy^2 - 5y^2 + 5)^2 y^2 + (x^4 + 24y^4 + 10x^3 - 10x - 25)^2)}{100} \right)$$

A view of the non-vanishing components of γ w.r.t. coordinates (x, y, ϕ) :

```
In [12]: gam.display_comp()
```

Out[12]:

$$\gamma_{xx} = \frac{96(x^3 - xy^2 - 5y^2 + 5)^2 y^2 + (x^4 + 24y^4 + 10x^3 - 10x - 25)^2}{100(x^2 - y^2)^3(x^2 - 1)}$$

$$\gamma_{yy} = -\frac{96(x^3 - xy^2 - 5y^2 + 5)^2 y^2 + (x^4 + 24y^4 + 10x^3 - 10x - 25)^2}{100(x^2 - y^2)^3(y^2 - 1)}$$

$$\gamma_{\phi\phi} = \frac{\left((96(x^3 - xy^2 - 5y^2 + 5)^2 y^2 + (x^4 + 24y^4 + 10x^3 - 10x - 25)^2)^2 (x^2 - 1) + 9600 \right) \left((24(y^2 - 1)^3(x + 5) + (2x^4 - (x^2 + 3)(y^2 - 1) - 2)(x^2 - 1) + 4(x^2 - y^2)x^2 + (x^2 - 1)(y^2 - 1))(x^2 - 1) \right)}{100 \left((96(x^2 - y^2)^2(x^2 - 1)(y^2 - 1) + ((x^2 - 1)^2 + 24(y^2 - 1)^2)^2 \right) (96(x^3 - xy^2 - 5y^2 + 5)^2 y^2 + (x^4 + 24y^4 + 10x^3 - 10x - 25)^2)}$$

Lapse function and shift vector

```
In [13]: N2 = AA2/BB2 - 2*m*q*CC2*(y^2-1)/BB2*(2*m*q*CC2*(y^2-1)
          / (BB2*(m^2*(y^2-1)*(p^2*BB2^2*(x^2-1)
          +4*q^2*d^2*CC2^2*(y^2-1))/(AA2*BB2*d^2)))
N2.simplify_full()
```

```
Out[13]: x10 + 20x9 + 576(x2 - 1)y8 + 99x8 - 40x7 + 96
          (x4 + 10x3 + 24x2 - 10x - 25)y6 - 350x6 - 480x5 - 48
          (3x6 + 10x5 - 3x4 + 20x3 + 125x2 - 30x - 125)y4 + 350x4 + 1000x3
          + 96(x8 - x6 + 10x5 - 10x3 + 25x2 - 25)y2 + 525x2 - 500x - 625
          -----
          x10 + 40x9 + 576(x2 - 1)y8 + 699x8 + 7920x7 + 96
          (x4 + 20x3 + 174x2 + 980x + 2425)y6 + 39450x6 - 960x5 - 48
          (3x6 + 20x5 - 3x4 + 40x3 + 925x2 + 5940x + 14675)y4 - 39450x4 - 6000
          x3 + 96(x8 - x6 + 20x5 - 20x3 + 375x2 + 3000x + 7425)y2 - 9675x2
          - 97000x - 240625
```

```
In [14]: N = Sig.scalar_field(sqrt(N2.simplify_full()), name='N')
print(N)
N.display()
```

Scalar field N on the 3-dimensional differentiable manifold Sigma

```
Out[14]: N : Σ → ℝ
          (x, y, φ) ↦ √ [ (x10+20x9+576(x2-1)y8+99x8-40x7+96(x4+10x3+24x2-10x-25)y6-350x6-480x5-48(3x6+10x5-3x4+20x3+125x2-30x-125)y4+350x4+1000x3+96(x8-x6+10x5-10x3+25x2-25)y2+525x2-500x-625) / (x10+40x9+576(x2-1)y8+699x8+7920x7+96(x4+20x3+174x2+980x+2425)y6+39450x6-960x5-48(3x6+20x5-3x4+40x3+925x2+5940x+14675)y4-39450x4-6000x3+96(x8-x6+20x5-20x3+375x2+3000x+7425)y2-9675x2-97000x-240625) ]
```

The coordinate expression of the scalar field N:

```
In [15]: N.expr()
```

```
Out[15]: √ [ x10 + 20x9 + 576(x2 - 1)y8 + 99x8 - 40x7 + 96
          (x4 + 10x3 + 24x2 - 10x - 25)y6 - 350x6 - 480x5 - 48
          (3x6 + 10x5 - 3x4 + 20x3 + 125x2 - 30x - 125)y4 + 350x4 + 1000x3
          + 96(x8 - x6 + 10x5 - 10x3 + 25x2 - 25)y2 + 525x2 - 500x - 625
          -----
          x10 + 40x9 + 576(x2 - 1)y8 + 699x8 + 7920x7 + 96
          (x4 + 20x3 + 174x2 + 980x + 2425)y6 + 39450x6 - 960x5 - 48
          (3x6 + 20x5 - 3x4 + 40x3 + 925x2 + 5940x + 14675)y4 - 39450x4 - 6000
          x3 + 96(x8 - x6 + 20x5 - 20x3 + 375x2 + 3000x + 7425)y2 - 9675x2
          - 97000x - 240625 ]
```

```
In [16]: b3 = 2*m*q*CC2*(y^2-1)/(BB2*(m^2*(y^2-1)*(p^2*BB2^2*(x^2-1)
          +4*q^2*d^2*CC2^2*(y^2-1))/(AA2*BB2*d^2))
b = Sig.vector_field('beta', latex_name=r'\beta')
b[3] = b3.simplify_full()
# unset components are zero
b.display_comp(only_nonzero=False)
```

```
Out[16]:  $\beta^x = 0$ 
 $\beta^y = 0$ 
```

$$\beta^\phi = -\frac{400(2\sqrt{6}x^7+24(\sqrt{6}x+5\sqrt{6})y^6+20\sqrt{6}x^6-\sqrt{6}x^5-72(\sqrt{6}x+5\sqrt{6})y^4-25\sqrt{6}x^4 - (\sqrt{6}x^5+15\sqrt{6}x^4+2\sqrt{6}x^3-10\sqrt{6}x^2-75\sqrt{6}x-365\sqrt{6})y^2+10\sqrt{6}x^2-25\sqrt{6}x-125\sqrt{6})}{x^{10}+40x^9+576(x^2-1)y^8+699x^8+7920x^7+96(x^4+20x^3+174x^2+980x+2425)y^6+39450x^6-960(3x^6+20x^5-3x^4+40x^3+925x^2+5940x+14675)y^4-39450x^4-6000x^3+96(x^8-x^6+20x^5-20x^3+375x^2+3000x+7425)y^2-9675x^2-97000x-240625}$$

Extrinsic curvature of Σ

We use the formula

$$K_{ij} = \frac{1}{2N} \mathcal{L}_\beta \gamma_{ij},$$

which is valid for any stationary spacetime:

```
In [17]: K = gam.lie_derivative(b) / (2*N)
K.set_name('K')
print(K)
```

Field of symmetric bilinear forms K on the 3-dimensional differentiable manifold Sigma

The component $K_{13} = K_{x\phi}$:

In [18]: K[1,3]

Out[18]:

2

$$\left(
\begin{aligned}
& 6\sqrt{3}\sqrt{2}x^{16} - 13824(\sqrt{3}\sqrt{2}x^2 + 10\sqrt{3}\sqrt{2}x + \sqrt{3}\sqrt{2})y^{16} + 240\sqrt{3}\sqrt{2}x^{15} \\
& + 3793\sqrt{3}\sqrt{2}x^{14} - 6912 \\
& (\sqrt{3}\sqrt{2}x^4 + 20\sqrt{3}\sqrt{2}x^3 + 150\sqrt{3}\sqrt{2}x^2 + 500\sqrt{3}\sqrt{2}x + 817\sqrt{3}\sqrt{2})y^{14} \\
& + 27650\sqrt{3}\sqrt{2}x^{13} + 72403\sqrt{3}\sqrt{2}x^{12} + 576 \\
& (27\sqrt{3}\sqrt{2}x^6 + 310\sqrt{3}\sqrt{2}x^5 + 1033\sqrt{3}\sqrt{2}x^4 + 1060\sqrt{3}\sqrt{2}x^3 + 10493\sqrt{3}\sqrt{2}x^2 \\
& + 44870\sqrt{3}\sqrt{2}x + 69503\sqrt{3}\sqrt{2}) \\
& - 81820\sqrt{3}\sqrt{2}x^{11} - 374975\sqrt{3}\sqrt{2}x^{10} - 96 \\
& (109\sqrt{3}\sqrt{2}x^8 + 520\sqrt{3}\sqrt{2}x^7 + 1504\sqrt{3}\sqrt{2}x^6 + 19360\sqrt{3}\sqrt{2}x^5 + 92770\sqrt{3}\sqrt{2}x^4 \\
& + 157960\sqrt{3}\sqrt{2}x^3 + 148264\sqrt{3}\sqrt{2}x^2 + 731920\sqrt{3}\sqrt{2}x + 1256425\sqrt{3}\sqrt{2}) \\
& - 313810\sqrt{3}\sqrt{2}x^9 + 669975\sqrt{3}\sqrt{2}x^8 + 24 \\
& (9\sqrt{3}\sqrt{2}x^{10} + 250\sqrt{3}\sqrt{2}x^9 + 6873\sqrt{3}\sqrt{2}x^8 + 40920\sqrt{3}\sqrt{2}x^7 + 63402\sqrt{3}\sqrt{2}x^6 \\
& + 146220\sqrt{3}\sqrt{2}x^5 + 1047426\sqrt{3}\sqrt{2}x^4 + 2249400\sqrt{3}\sqrt{2}x^3 + 876525\sqrt{3}\sqrt{2}x^2 \\
& + 4308810\sqrt{3}\sqrt{2}x + 8401925\sqrt{3}\sqrt{2}) \\
& + 1617000\sqrt{3}\sqrt{2}x^7 + 999675\sqrt{3}\sqrt{2}x^6 + 96 \\
& (20\sqrt{3}\sqrt{2}x^{11} - 179\sqrt{3}\sqrt{2}x^{10} - 50\sqrt{3}\sqrt{2}x^9 - 2897\sqrt{3}\sqrt{2}x^8 - 28400\sqrt{3}\sqrt{2}x^7 \\
& - 57446\sqrt{3}\sqrt{2}x^6 - 9020\sqrt{3}\sqrt{2}x^5 - 237650\sqrt{3}\sqrt{2}x^4 - 731060\sqrt{3}\sqrt{2}x^3 \\
& - 267175\sqrt{3}\sqrt{2}x^2 - 1037250\sqrt{3}\sqrt{2}x - 2111325\sqrt{3}\sqrt{2}) \\
& - 2277250\sqrt{3}\sqrt{2}x^5 - 4979375\sqrt{3}\sqrt{2}x^4 \\
& - (187\sqrt{3}\sqrt{2}x^{14} + 3590\sqrt{3}\sqrt{2}x^{13} - 5207\sqrt{3}\sqrt{2}x^{12} - 73540\sqrt{3}\sqrt{2}x^{11} - 45465\sqrt{3}\sqrt{2}x^{10} \\
& - 1150150\sqrt{3}\sqrt{2}x^9 + 199401\sqrt{3}\sqrt{2}x^8 - 1059000\sqrt{3}\sqrt{2}x^7 \\
& - 7811175\sqrt{3}\sqrt{2}x^6 + 2899610\sqrt{3}\sqrt{2}x^5 + 1675075\sqrt{3}\sqrt{2}x^4 - 32834500\sqrt{3}\sqrt{2}x^3 \\
& - 24681575\sqrt{3}\sqrt{2}x^2 - 69684250\sqrt{3}\sqrt{2}x - 122823125\sqrt{3}\sqrt{2}) \\
& - 4037500\sqrt{3}\sqrt{2}x^3 + 3461875\sqrt{3}\sqrt{2}x^2 - 6 \\
& (\sqrt{3}\sqrt{2}x^{16} + 40\sqrt{3}\sqrt{2}x^{15} + 601\sqrt{3}\sqrt{2}x^{14} + 4010\sqrt{3}\sqrt{2}x^{13} + 12935\sqrt{3}\sqrt{2}x^{12}) \\
& - 1060\sqrt{3}\sqrt{2}x^{11} + 10449\sqrt{3}\sqrt{2}x^{10} + 139590\sqrt{3}\sqrt{2}x^9 + 57825\sqrt{3}\sqrt{2}x^8 \\
& + 146960\sqrt{3}\sqrt{2}x^7 + 781475\sqrt{3}\sqrt{2}x^6 - 702250\sqrt{3}\sqrt{2}x^5 - 2108075\sqrt{3}\sqrt{2}x^4 \\
& - 348500\sqrt{3}\sqrt{2}x^3 + 2381875\sqrt{3}\sqrt{2}x^2 + 5456250\sqrt{3}\sqrt{2}x + 6941250\sqrt{3}\sqrt{2}) \\
& + 7231250\sqrt{3}\sqrt{2}x + 6109375\sqrt{3}\sqrt{2}
\end{aligned}
\right)$$

The type-(1,1) tensor K^\sharp of components $K^i_j = \gamma^{ik} K_{kj}$:

```
In [19]: Ku = K.up(gam, 0)
print(Ku)
```

Tensor field of type (1,1) on the 3-dimensional differentiable manifold Sigma

We may check that the hypersurface Σ is maximal, i.e. that $K^k_k = 0$:

```
In [20]: trK = Ku.trace()
trK
```

```
Out[20]: 0
```

Connection and curvature

Let us call D the Levi-Civita connection associated with γ :

```
In [21]: D = gam.connection(name='D')
print(D)
```

Levi-Civita connection D associated with the Riemannian metric gam on the 3-dimensional differentiable manifold Sigma

The Ricci tensor associated with γ :

```
In [22]: Ric = gam.ricci()
print(Ric)
```

Field of symmetric bilinear forms Ric(gam) on the 3-dimensional differentiable manifold Sigma

The scalar curvature $R = \gamma^{ij} R_{ij}$:

```
In [23]: R = gam.ricci_scalar(name='R')
print(R)
```

Scalar field R on the 3-dimensional differentiable manifold Sigma

The coordinate expression of the Ricci scalar is huge:

In [24]: R.expr()

Out[24]: 480000

$$\begin{aligned}
& 36x^{38} + 2880x^{37} - 191102976(3x^4 + 20x^3 - 6x^2 - 60x - 101)y^{36} \\
& + 103116x^{36} + 2152440x^{35} + 382205952 \\
& (5x^6 + 45x^5 + 195x^4 + 1210x^3 + 3555x^2 + 4065x - 291)y^{34} + 28527685 \\
& x^{34} + 243524500x^{33} - 63700992 \\
& (45x^8 + 605x^7 + 4434x^6 + 21165x^5 + 36233x^4 - 28965x^3 - 244428x^2y^{32} \\
& - 457885x - 503220) \\
& + 1269998358x^{32} + 3199445660x^{31} + 10616832 \\
& (279x^{10} + 4375x^9 + 30087x^8 + 71400x^7 - 238782x^6 - 1698210x^5 - y^{30} \\
& - 6419238x^4 - 18093840x^3 - 36663561x^2 - 47421645x - 39429937) \\
& - 2269601041x^{30} - 34623715080x^{29} - 331776 \\
& (5729x^{12} + 49180x^{11} - 74502x^{10} - 3734980x^9 - 26371593x^8 - 109948680y^{28} \\
& x^7 - 472804404x^6 - 1606840680x^5 - 3848737185x^4 - 6740837780x^3 \\
& - 9099798310x^2 - 9444811860x - 7651340375) \\
& - 59087224000x^{28} + 65688034640x^{27} - 110592 \\
& (2649x^{14} + 238580x^{13} + 3051495x^{12} + 19491840x^{11} + 91830265x^{10} - y^{26} \\
& + 456692580x^9 + 2183925951x^8 + 7465185120x^7 + 19582237971x^6 \\
& + 43189401660x^5 + 79002052285x^4 + 114585899040x^3 + 122428856475x^2 \\
& + 101684757740x + 84896596125) \\
& + 384825320925x^{26} + 395426661500x^{25} + 9216 \\
& (181479x^{16} + 4048360x^{15} + 34156308x^{14} + 188537160x^{13} + 1036205812 - y^{24}
\end{aligned}$$

3+1 Einstein equations

Let us check that the vacuum 3+1 Einstein equations are satisfied.

We start by the constraint equations:

Hamiltonian constraint

Let us first evaluate the term $K_{ij}K^{ij}$:

```
In [25]: Kuu = Ku.up(gam, 1)
trKK = K['_ij']*Kuu['^ij']
print(trKK)
```

Scalar field on the 3-dimensional differentiable manifold Sigma

The vacuum Hamiltonian constraint equation is

$$R + K^2 - K_{ij}K^{ij} = 0$$

```
In [26]: Ham = R + trK^2 - trKK
print(Ham)
Ham.display()
```

Scalar field zero on the 3-dimensional differentiable manifold Sigma

```
Out[26]: 0: Σ      → ℝ
        (x, y, φ) ↦ 0
```

Hence the Hamiltonian constraint is satisfied.

Momentum constraint

In vacuum, the momentum constraint is

$$D_j K^j_i - D_i K = 0$$

```
In [27]: mom = D(Ku).trace(0,2) - D(trK)
print(mom)
mom.display()
```

1-form on the 3-dimensional differentiable manifold Sigma

```
Out[27]: 0
```

Hence the momentum constraint is satisfied.

Dynamical Einstein equations

Let us first evaluate the symmetric bilinear form $k_{ij} := K_{ik}K^k_j$:

```
In [28]: KK = K['_ik']*Ku['^k_j']
print(KK)
```

Tensor field of type (0,2) on the 3-dimensional differentiable manifold Sigma

```
In [29]: KK1 = KK.symmetrize()
         KK == KK1
```

Out[29]: True

```
In [30]: KK = KK1
         print(KK)
```

Field of symmetric bilinear forms on the 3-dimensional differentiable manifold Sigma

In vacuum and for stationary spacetimes, the dynamical Einstein equations are

$$\mathcal{L}_\beta K_{ij} - D_i D_j N + N (R_{ij} + K K_{ij} - 2 K_{ik} K_j^k) = 0$$

```
In [31]: dyn = K.lie_derivative(b) - D(D(N)) + N*( Ric + trK*K - 2*KK )
         print(dyn)
         dyn.display()
```

Tensor field of type (0,2) on the 3-dimensional differentiable manifold Sigma

Out[31]: 0

Hence the dynamical Einstein equations are satisfied.

Finally we have checked that all the 3+1 Einstein equations are satisfied by the $\delta = 2$ Tomimatsu-Sato solution.